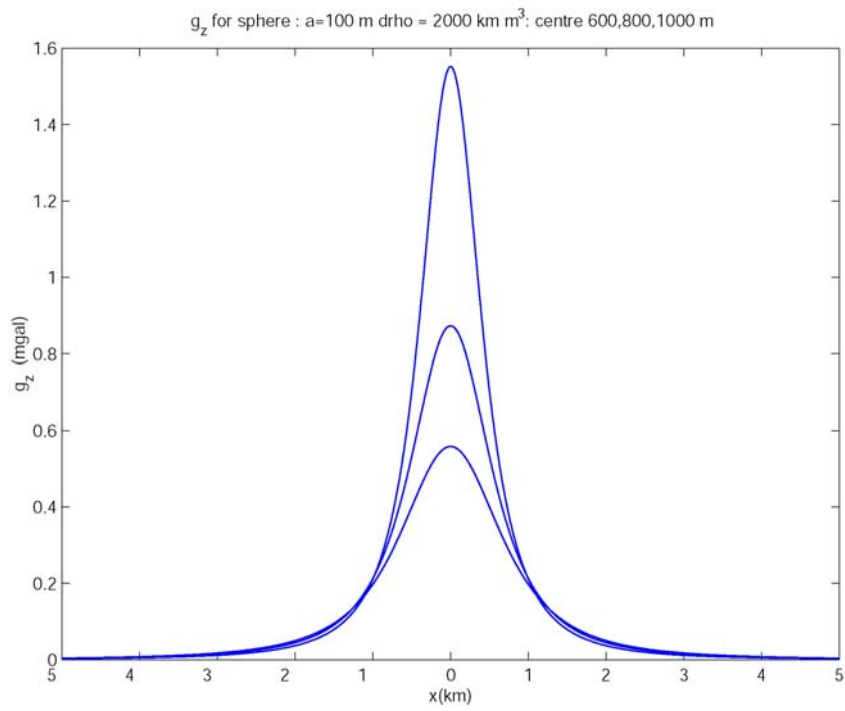
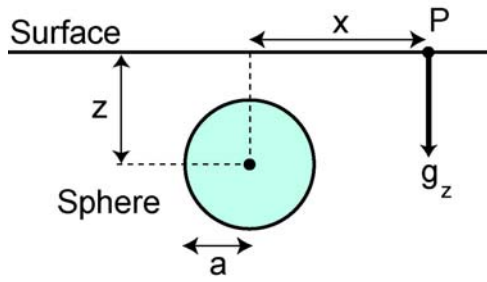


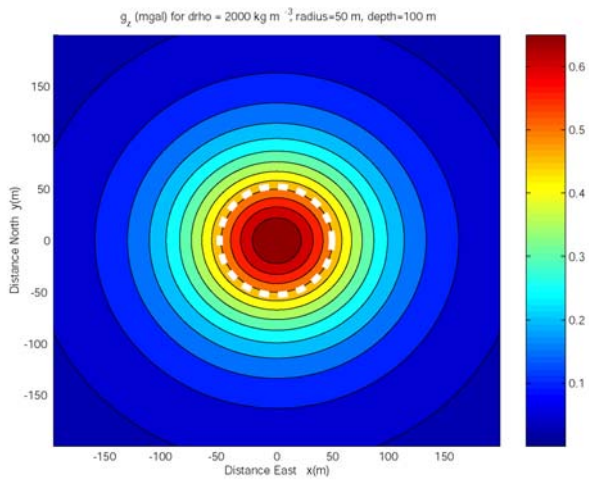
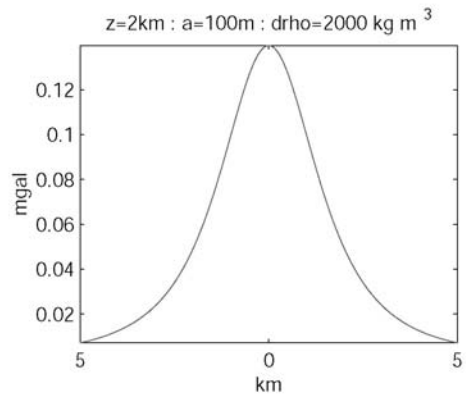
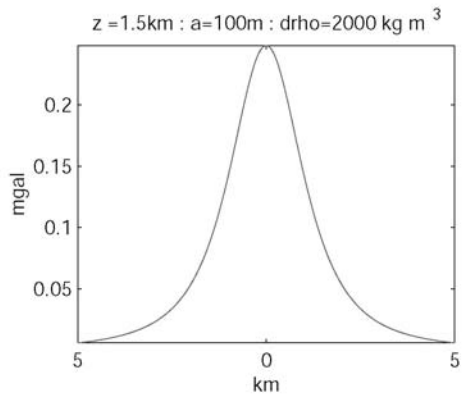
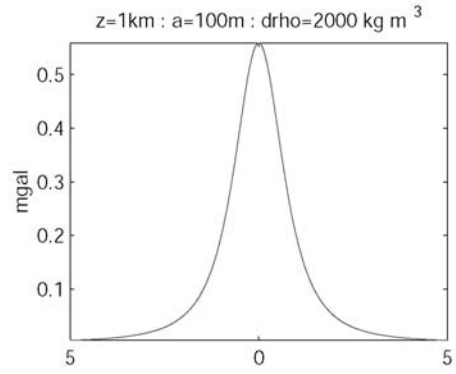
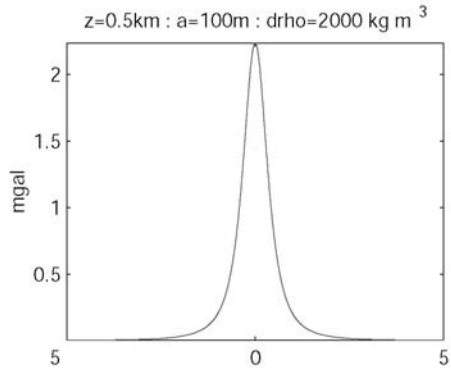
## **Geophysics 325**

**B3.1 Gravitational acceleration due to a thin rod**

**B3.2 Gravitational acceleration due to a thin sheet**

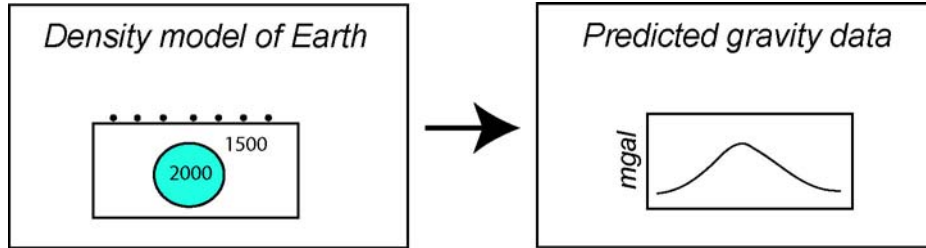
### B3.3 Gravitational acceleration due to a buried sphere





## Note on Forward and inverse problems in Geophysics

### Forward problems

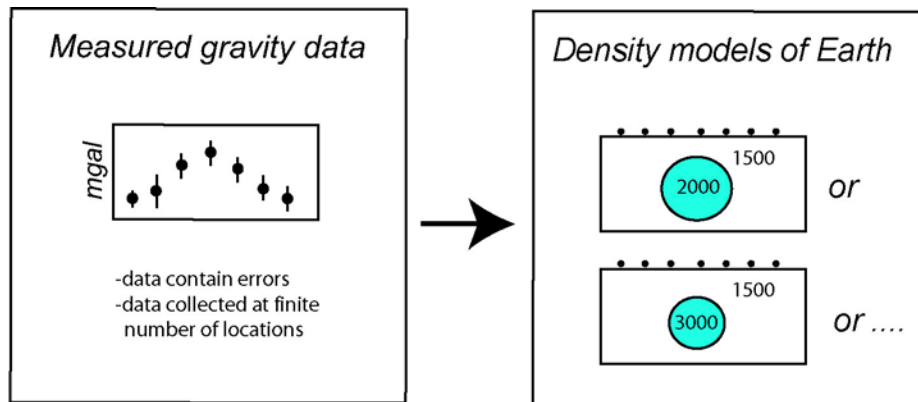


-start with a model of the Earth and compute predicted data

-unique solution

-calculating the predicted data can be a simple equation or a complex computer algorithm

### Inverse problems



-start with measured data and compute a model of the Earth (density, velocity etc)

-non-unique with many solutions

-solving the inverse problem can involve a trial-and-error approach or an automated inversion algorithm

Non-uniqueness can arise from two distinct factors

### (1) The basic physics

Example : In gravity exploration, we can only determine the excess mass of a buried sphere. We cannot determine the density and radius that combine to give this value of excess mass.

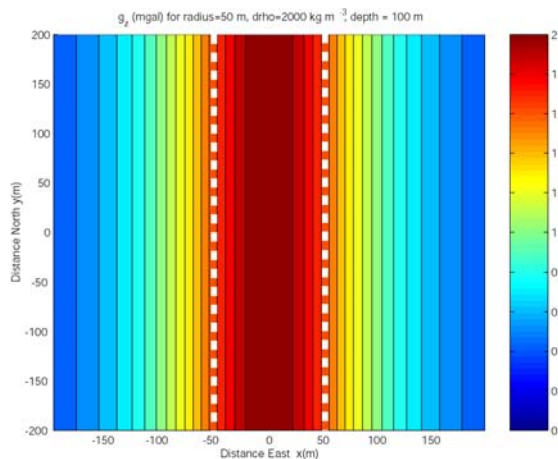
This type of non-uniqueness cannot be overcome, not even with expensive computer packages. However, additional (independent) data can be used to address non-uniqueness. For example, if we have density measurements of the target, we could determine the radius of the sphere.

### (2) Noise in data

Example: In gravity exploration noise in the data will introduce errors when we determine the half width ( $x_{1/2}$ ). Errors in the half width will introduce errors into estimates of the depth of sphere.

This type of non-uniqueness can be overcome by improving data quality and quantity.

## B3.4 Gravitational acceleration due to a buried cylinder



### B3.5 Gravitational acceleration of a 2-D polygon

Integration is needed to compute the gravitational acceleration of an arbitrary shape. If the density is a function of  $x, y$  and  $z$ , then the approach in B2.1 must be used (integrate over each cubic cell). However, if a body has a uniform density contrast,  $\Delta\rho$ , then this can be reduced to a surface integral. In the case of a 2-D structure, the body will be a prism that is invariant in the  $y$ -direction. Talwani (1959) showed that the gravity acceleration can be computed from a line integral around the perimeter.

$$g_z = 2G\Delta\rho \oint \tan^{-1}\left(\frac{x}{z}\right) dz$$

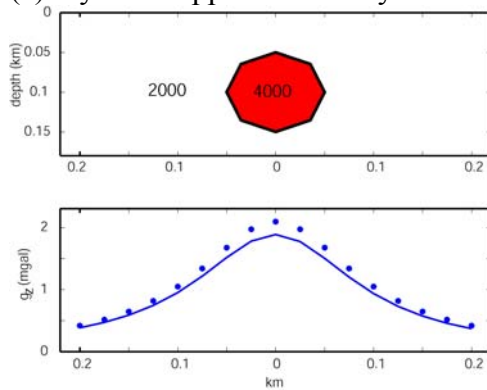
This algorithm was implemented in a MATLAB script `g_polygon_N.m`

#### Example 1 Horizontal cylinder

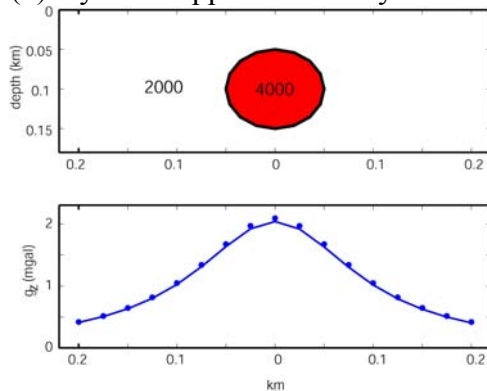
Density contrast =  $2000 \text{ kg m}^{-3}$

Before using any software, we need to test algorithm against a known result. Be suspicious of all software.

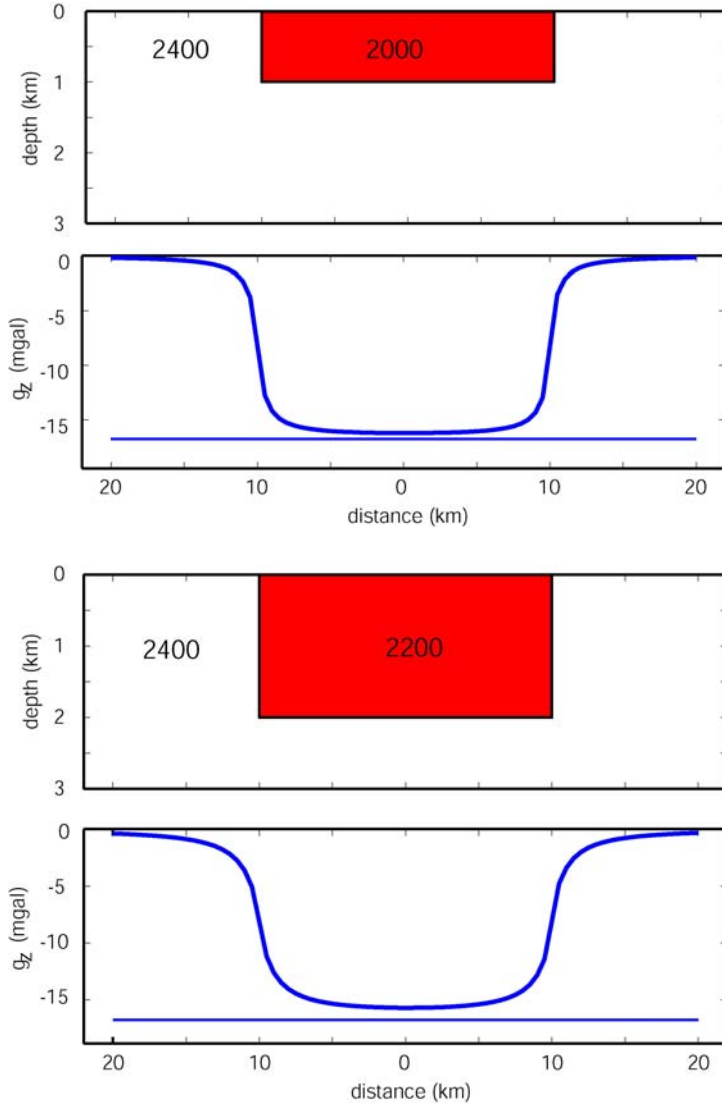
(a) Cylinder approximated by a 8-sided polygon



(b) Cylinder approximated by a 16-sided polygon



## Example 2 : Low density basin



Horizontal line shows the value of  $g_z$  for an infinite slab, computed with the result derived in B3.2

How far from the edges do we need to be to use the formula for an infinite slab?

The 2-D line integral approach will be used later in the class to interpret some real gravity anomaly data collected in a number of locations.