

SOLUTION

$$1(a) \quad g_S = \frac{GM_S}{R^2} = 5.89 \times 10^{-3} \text{ ms}^{-2} = 589 \text{ mgal}$$

$$(b) \quad g_M = \frac{GM_M}{R^2} = 3.33 \times 10^{-5} \text{ ms}^{-2} = 3.3 \text{ mgals}$$

note units for  $R \Rightarrow \text{km}$

(c) Tides due to gradient of gravity. Thus since  $g \propto 1/r^2$  the gradient is highest for the moon.

Keyword: spatial gradient. Also note  $g_S > g_M$

(d) Length of day observations

$\Rightarrow$  accurate astronomy

$\Rightarrow$  timing of eclipses

$\Rightarrow$  coral growth rings (daily)

$\Rightarrow$  tidal rhythmites

see notes

(e) See notes. Keywords - tidal friction, slows earth

$\Rightarrow$  conserve angular momentum  
requires moon to orbit faster

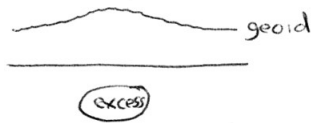
$\Rightarrow$  higher orbit.

Q2 (a) Distribution  $g_P > g_E$   
mass re-distribution  $g_P < g_E$   
rotation  $g_P > g_E$

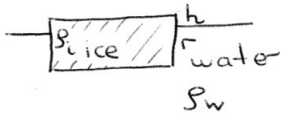
(b)  $\rightarrow$  sea level measured by radar from a satellite  
Sea bunches up over an excess mass

$\rightarrow$  monitor orbit of satellite. Either track from ground, or with a pair of satellites (GRACE)

(c) Geoid - gravitational, equipotential at mean sea-level



Q3



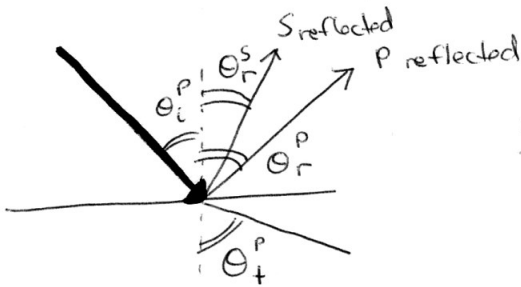
$$r = \frac{\rho_i}{\rho_w - \rho_i} h$$

$$r + h = 400$$

$$h = 42.7 \text{ m}$$

Q4

⇒ no S-wave in outer core. It is liquid (notes)



$$\frac{\sin \theta_i^P}{v_p} = 0.04412$$

P-wave reflected

$$0.04412 = \frac{\sin \theta_r^P}{13}$$

$$\theta_r^P = 35^\circ$$

S-wave reflected

$$0.04412 = \frac{\sin \theta_r^S}{7}$$

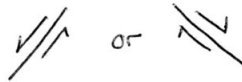
$$\theta_r^S = 18.0^\circ$$

P-wave transmitted

$$0.04412 = \frac{\sin \theta_t^P}{8}$$

$$\theta_t^P = 20.7^\circ$$

Q5(a)

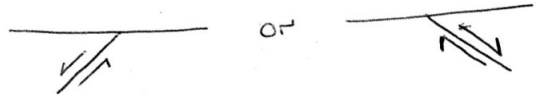


• two possible strike-slip faults



map

normal fault



• two possible normal faults

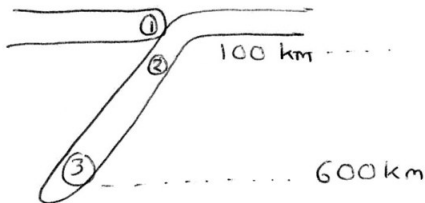


⇒ two possible orientations  
⇒ mixture of strike-slip and normal fault;



⇒ more small earthquakes than large earthquakes

(c)



- ① MEGATHRUST
- ② INTERMEDIATE
- ③ DEEP