

A nonlinear stability theorem for baroclinic quasigeostrophic flow

Gordon E. Swaters

Department of Earth, Atmospheric, and Planetary Sciences, Massachusetts Institute of Technology,
Cambridge, Massachusetts 02139

(Received 23 August 1985; accepted 22 October 1985)

The baroclinic quasigeostrophic equations describe the essential dynamics of large-scale, low-frequency atmospheric and oceanic flow. A nonlinear stability theorem is given based on a convexity argument of Arnold [Am. Math. Soc. Transl. **19**, 267 (1969)], complementing a linear analysis by Blumen [J. Atmos. Sci. **25**, 929 (1968)]. An *a priori* estimate bounding the growth of perturbations is derived.

Recently, several studies have attempted to establish the nonlinear stability of various planetary flows.¹⁻⁵ All of these analyses have been based on deriving sufficient conditions for the positive definiteness of the second variation of a relevant constrained Hamiltonian describing the basic flow. However, it is known⁶⁻⁸ that this method fails to establish nonlinear stability for infinite-dimensional dynamical systems. Rigorous nonlinear stability theorems require certain convexity hypotheses on the constrained Hamiltonian.⁹⁻¹¹ For flows of relevance to geophysical fluid dynamics, correct nonlinear stability proofs have been given for the stratified Euler equations,¹² compressible barotropic flows,¹³ circular vortex patches,¹⁴ and multilayer quasigeostrophic flows.¹⁵ The principle purpose of this letter is to provide a rigorous nonlinear stability theorem for continuously stratified quasigeostrophic flow including the effects of topography in a bounded or unbounded horizontal domain. (See note added in proof.) These equations describe the essential dynamics for large-scale, low-frequency atmospheric and oceanic motions.^{16,17}

The nondimensional quasigeostrophic equations for a vertically stratified fluid on the beta plane in the absence of heating or dissipation are

$$[\partial_t + J(p, \cdot)] [\Delta p + \bar{\rho}^{-1} (\bar{\rho} S^{-1} p_z)_z + \beta y] = 0, \quad (1)$$

with

$$[\partial_t + J(p, \cdot)] [p_z - S(z)h(x, y)] = 0 \quad \text{on } z = 0, \quad (2)$$

$$[\partial_t + J(p, \cdot)] p_z = 0 \quad \text{on } z = 1, \quad (3)$$

where $p = C_0$ on the smooth horizontal boundary D , or $\nabla p \rightarrow 0$ as $(x^2 + y^2) \rightarrow \infty$ if the fluid is horizontally unbounded. The symbols are standard¹⁶⁻¹⁸; however, we point out here that $h(x, y)$ is the topography, p is the geostrophic pressure, the two-dimensional geostrophic velocity field $(v, v) = (-p_y, p_x)$, $J(\cdot, \cdot) = \partial(\cdot, \cdot)/\partial(x, y)$, and derivatives are denoted by subscripts. Note that $\nabla = (\partial_x, \partial_y)$ and $\Delta = \nabla^2$.

The nonlinear stability analysis presented here is based on a Hamiltonian formulation of (1)–(3) given in a previous linear analysis by Blumen,¹⁹ modified to include the effects of topography and a smooth arbitrary horizontal boundary. The analysis begins by noting that the functions

$$E_p = \frac{1}{2} \iiint \bar{\rho} [\nabla p \cdot \nabla p + S^{-1} (p_z)^2] dV,$$

$$F_1 = \iiint \bar{\rho} \Phi_1(z; q) dV,$$

$$F_2 = - \iint [\bar{\rho} S^{-1} \Phi_2(p_z)]_{z=1} dA,$$

$$F_3 = \iint [\bar{\rho} S^{-1} \Phi_3(p_z - Sh)]_{z=0} dA,$$

$$F_4 = - \int \lambda \bar{\rho} \oint_D \mathbf{n} \cdot \nabla p ds dz,$$

are conserved by (1)–(3),¹⁶ where Φ_1 , Φ_2 , and Φ_3 are smooth functions of their arguments; the quasigeostrophic potential vorticity is given by $q = \Delta p + \bar{\rho}^{-1} \times (\bar{\rho} S^{-1} p_z)_z + \beta y$, \mathbf{n} is the outward unit normal on D , $dV = dx dy dz$, and $dA = dx dy$.

The Hamiltonian is given by

$$H = E_p + F_1 + F_2 + F_3 + F_4, \quad (4)$$

and an equilibrium solution of (1)–(3) is denoted as p^s . The Hamiltonian can be chosen as (4) since F_1 , F_2 , and F_3 are Casimir functions with respect to the mass-constrained Hamiltonian $E_p + F_4$ and Poisson bracket given in Ref. 20. The derivative of H evaluation at p^s is given by (after integration by parts)

$$\begin{aligned} DH(p^s) \delta p = & \iiint \bar{\rho} \delta q [\Phi_1'(q^s) - p^s] dV \\ & + \iint [\bar{\rho} S^{-1} \delta p_z [p^s - \Phi_2'(p_z^s)]]_{z=1} dA \\ & + \iint \{ \bar{\rho} S^{-1} \delta p_z \\ & \times [\Phi_3'(p_z^s - Sh) - p^s] \}_{z=0} dA \\ & + \int \bar{\rho} \oint_D (C_0 - \lambda) \mathbf{n} \cdot \nabla \delta p ds dz, \end{aligned}$$

where $\delta p = p - p^s$, $\delta q = q - q^s$, and $(\cdot)'$ means differentiation with respect to the argument. The stationary solution p^s is a critical point of H when

$$\Phi_1'(q^s) = p^s, \quad (5)$$

$$\Phi_2'(p_z^s) = p^s \quad \text{on } z = 1, \quad (6)$$

$$\Phi_3'(p_z^s - Sh) = p^s \quad \text{on } z = 0, \quad (7)$$

and when $\lambda = C_0$. Equations (5)–(7) serve to define the functions Φ_1 , Φ_2 , and Φ_3 given p^s .

Nonlinear stability is proved by assuming the convexity hypotheses

$$0 < \alpha_1 < \Phi_1'' < \alpha_2 < \infty, \quad (8)$$

$$0 < \mu_1 < -\Phi_2'' < \mu_2 < \infty, \quad (9)$$

$$0 < \gamma_1 < \Phi_3'' < \gamma_2 < \infty, \quad (10)$$

and examining the conserved functional

$$\hat{H}(\delta p) \equiv H(p^s + \delta p) - H(p^s) - DH(p^s) \delta p. \quad (11)$$

Here \hat{H} is conserved since H is conserved, and $DH(p^s)\delta p = 0$ for p^s satisfying (5)–(7). The conditions (8)–(10) are assumed to hold for all arguments and at all times for which smooth solutions exist to (1)–(3). When attention is restricted to a basic state corresponding to a zonal flow [i.e., $p^s = p^s(y, z)$], the convexity assumptions (8)–(10) can be shown to be equivalent to Pedlosky's sufficient linear stability conditions for a zonal flow.²¹

It follows from (11), exploiting the convexity hypotheses (8)–(10), that

$$\begin{aligned} 2E_{\delta p} + \alpha_1 \iint \int \bar{\rho}(\delta q)^2 dV &+ \mu_1 \iint [\bar{\rho}S^{-1}(\delta p_z)^2]_{z=1} dA \\ &+ \gamma_1 \iint [\bar{\rho}S^{-1}(\delta p_z)^2]_{z=0} dA < 2\hat{H}(\delta p) < 2E_{\delta p} \\ &+ \alpha_2 \iint \int \bar{\rho}(\delta q)^2 dV + \mu_2 \iint [\bar{\rho}S^{-1}(\delta p_z)^2]_{z=1} dA \\ &+ \gamma_2 \iint [\bar{\rho}S^{-1}(\delta p_z)^2]_{z=0} dA. \end{aligned}$$

Since $\hat{H}(\delta p) = \hat{H}(\delta p_0)$, where $\delta p_0 \equiv \delta p(t=0)$, it follows that

$$\begin{aligned} 2E_{\delta p} + \alpha_1 \iint \int \bar{\rho}(\delta q)^2 dV &+ \mu_1 \iint [\bar{\rho}S^{-1}(\delta p_z)^2]_{z=1} dA \\ &+ \gamma_1 \iint [\bar{\rho}S^{-1}(\delta p_z)^2]_{z=0} dA \\ &< 2E_{\delta p_0} + \alpha_2 \iint \int \bar{\rho}(\delta q_0)^2 dV \\ &+ \mu_2 \iint [\bar{\rho}S^{-1}(\delta p_{0z})^2]_{z=1} dA \\ &+ \gamma_2 \iint [\bar{\rho}S^{-1}(\delta p_{0z})^2]_{z=0} dA, \quad (12) \end{aligned}$$

which establishes nonlinear stability. The *a priori* estimate (12) implies Lyapunov stability of smooth solutions to (1)–(3) and is explicitly independent of the topography $h(x, y)$.²²

The existence of classical solutions of (1)–(3) in a horizontally periodic domain has been proved only up to a finite time, which is inversely proportional to the norms of p_{0z} on $z=0$ and $z=1$, and q_0 (Ref. 23). Stability can also be proved when $\Phi_1'' < 0$, and $\Phi_2'' > 0$ and $\Phi_3'' < 0$ by considering $-\hat{H}(\delta p)$ and requiring sufficiently large $\min(-\Phi_1'')$, $\min(\Phi_2'')$, and $\min(\Phi_3'')$. The stability theorem can be generalized to include "islands" (i.e., non-simply-connected domains) by introducing other circulation functions similar to

F_4 . Since the quasigeostrophic equations are zonally (i.e., in the x direction) Galilean invariant, these results also apply to (zonally) steadily translating fluid motions. Therefore the stability theorem presented here is of importance for the nonlinear stability of solitary planetary waves²⁴ (provided Φ_1 , Φ_2 , and Φ_3 are sufficiently smooth). On this latter application, a more detailed analysis will be published elsewhere.

Note added in proof: The author has become aware of a similar stability analysis by M. E. McIntyre and T. G. Shepherd.²⁵

ACKNOWLEDGMENTS

The author thanks Dr. Michael E. McIntyre and Dr. Darryl D. Holm for helpful comments made on the manuscript.

This paper was written while the author was supported by a National Science Foundation Grant No. 8019260-OCE to Professor Glenn R. Flierl of the Massachusetts Institute of Technology.

- ¹R. Benzi, S. Pierini, and A. Vulpiani, *Geophys. Astrophys. Fluid Dyn.* **20**, 293 (1982).
- ²R. Petroni and A. Vulpiani, *Nuovo Cimento B* **78**, 1 (1983).
- ³S. Pierini and E. Salusti, *Nuovo Cimento B* **71**, 282 (1982).
- ⁴S. Pierini and A. Vulpiani, *J. Phys. A* **14**, L203 (1981).
- ⁵R. Purini and S. Salusti, *Geophys. Astrophys. Fluid Dyn.* **30**, 261 (1984).
- ⁶D. G. Ebin and J. E. Marsden, *Ann. Math.* **92**, 106 (1970).
- ⁷J. E. Marsden and R. Abraham, *Proc. Symp. Pure Math. (AMS)* **16**, 237 (1970).
- ⁸V. I. Arnold, *Mathematical Methods of Classical Mechanics* (Springer, Berlin, 1978), Appendices 2 and 5.
- ⁹V. I. Arnold, *Am. Math. Soc. Transl.* **19**, 267 (1969).
- ¹⁰J. M. Ball and J. E. Marsden, *Arch. Ration. Mech. Anal.* **86**, 251 (1984).
- ¹¹D. D. Holm, J. E. Marsden, T. Ratiu, and A. Weinstein, *Phys. Rep.* **123**, 32 (1985), Secs. 1–3.
- ¹²H. D. I. Abarbanel, D. D. Holm, J. E. Marsden, and T. Ratiu, *Phys. Rev. Lett.* **52**, 2352 (1984).
- ¹³D. D. Holm, J. E. Marsden, T. Ratiou, and A. Weinstein, *Phys. Lett.* **98A**, 15 (1983).
- ¹⁴Y. H. Wan and M. Pulvirenti, *Commun. Math. Phys.* **99**, 435 (1985).
- ¹⁵D. D. Holm, J. E. Marsden, T. Ratiu, and A. Weinstein, *Phys. Rep.* **123**, 32 (1985), Part I, Sec. 4.
- ¹⁶J. Pedlosky, *Geophysical Fluid Dynamics* (Springer, New York, 1979).
- ¹⁷P. H. LeBlond and L. A. Mysak, *Waves in the Ocean* (Elsevier, New York, 1978).
- ¹⁸We assume a stably stratified fluid so $S(z) > 0$. In the ocean the stratification parameter $S(z)$ is proportional to $-g\bar{\rho}^{-1}\bar{\rho}_z$ [one can set $\bar{\rho} = 1$ in (1)]. In the atmosphere $S(z) \sim g\Theta^{-1}\Theta_z$, where Θ is the ambient vertically stratified potential temperature.
- ¹⁹W. Blumen, *J. Atmos. Sci.* **25**, 929 (1968).
- ²⁰D. D. Holm, *Phys. Fluids* **29**, 7 (1986).
- ²¹In the absence of topography (i.e., $h=0$), this was demonstrated by Blumen.¹⁹ For the more general linear stability result, see Pedlosky (Chap. 7).¹⁶ Andrews [*Geophys. Astrophys. Fluid Dyn.* **28**, 243 (1984)] has shown that the only steady quasigeostrophic flows satisfying zonally symmetric boundary conditions for which (8)–(10) hold are zonally symmetric flows.
- ²²While $h(x, y)$ does not explicitly enter (12), it implicitly affects the *a priori* bound since γ_1 and γ_2 of (10) depend on $h(x, y)$.
- ²³A. F. Bennett and P. E. Kloeden, *Proc. R. Soc. Edinburgh Sect. A* **91**, 185–203 (1982).
- ²⁴P. Malanotte-Rizzoli, *Adv. Geophys.* **24**, 197 (1982).
- ²⁵M. E. McIntyre (private communication).

